

# Distance to which NEOcam could detect Chelyabinsk meteor that hit Russia Feb. 2013

$$r := \frac{20}{2} \cdot m \quad \text{radius of object, increased from 17 to 20 based on NOVA program}$$

$$T := 280 \cdot K \quad \text{from our L1 orbit, we would view the sunlit side of the object as it approached the earth. This estimate assumes rapid rotation, it would be hotter still otherwise.}$$

$$F_8 := 10^{-13} \cdot \frac{\text{erg}}{\text{cm}^2 \cdot \text{sec}} \cdot \frac{8 \cdot \mu\text{m}}{c} \quad F_8 = 2.669 \times 10^{-4} \text{ Jy} \quad \text{5 sigma flux limit NEOcam, 6-10 um band}$$

$$F = \frac{\pi \cdot r^2}{d^2} \cdot .95 \cdot B_v(v, T) \quad \text{Flux of object at distance d, 95% emissivity}$$

$$d := \sqrt{\frac{\pi \cdot r^2}{F_8} \cdot .95 \cdot B_v\left(\frac{c}{8 \cdot \mu\text{m}}, T\right)} \quad d = 0.079 \text{ AU} \quad d = 1.184 \times 10^7 \text{ km}$$

cf. Lagrange L1 point is 1.5 million km from earth, confirming that NEOcam could have detected this object well (~ 45 days) before it hit earth!

$$v := 18 \cdot \frac{\text{km}}{\text{sec}} \quad \text{velocity when it hit Earth}$$

$$m := \frac{4}{3} \cdot \pi \cdot r^3 \cdot 3 \cdot \frac{\text{gm}}{\text{cm}^3} \quad \text{mass of rocky object, density an estimate.} \quad m = 1.257 \times 10^7 \text{ kg}$$

$$\frac{1}{2} \cdot m \cdot v^2 = 0.497 \text{ Mton} \quad \text{*extremely* close to the 500 kton measurement of energy from infrasound.}$$

$$\begin{aligned} k &\equiv 1.38 \cdot 10^{-16} \cdot \text{erg} \cdot K^{-1} & M_{\text{sun}} &\equiv 2 \cdot 10^{33} \cdot \text{gm} & \text{AU} &\equiv 1.5 \cdot 10^{13} \cdot \text{cm} \\ R_{\text{sun}} &\equiv 7.0 \cdot 10^{10} \cdot \text{cm} & \text{arcsec} &\equiv \frac{\pi}{180 \cdot 3600} \cdot \text{rad} \\ h &\equiv 6.63 \cdot 10^{-27} \cdot \text{erg} \cdot \text{sec} & L_{\text{sun}} &\equiv 3.9 \cdot 10^{33} \cdot \frac{\text{erg}}{\text{sec}} & \text{pc} &\equiv \text{AU} \cdot \frac{\text{rad}}{\dots} \\ \text{star} &\equiv 1 \quad \text{dimensionless constant} & \text{sc} &\equiv 1.36 \cdot 10^6 \cdot \text{erg} \cdot \text{cm}^{-2} \cdot \text{sec}^{-1} & \text{pc} &\equiv 3.094 \times 10^{18} \text{ cm} \\ \text{mm} &\equiv 10^{-6} \cdot \text{m} \quad \text{W} = \text{watt} & \text{ly} &\equiv \text{c} \cdot \text{yr} \\ \text{Angstrom} &\equiv 10^{-8} \cdot \text{cm} & \text{nc} &\equiv 3.27 \text{ ly} \\ G &\equiv 6.67 \cdot 10^{-8} \cdot \text{erg} \cdot \frac{\text{cm}}{\text{gm}^2} & \sigma &\equiv 5.67 \cdot 10^{-5} \cdot \text{erg} \cdot \text{cm}^{-2} \cdot \text{sec}^{-1} \cdot K^{-4} \\ A &\equiv 6.022 \cdot 10^{23} & a &\equiv \frac{4 \cdot \sigma}{c} \end{aligned}$$

$$\begin{aligned} e_- &\equiv 1.602 \cdot 10^{-19} \cdot \text{coul} & e_- &\equiv 4.803 \times 10^{-10} \text{ esu} & \text{eV} &\equiv e_- \cdot \text{volt} \\ m_H &\equiv \frac{1 \cdot \text{gm}}{A} & m_H &\equiv 1.661 \times 10^{-24} \text{ gm} & \text{eV} &\equiv 1.602 \times 10^{-12} \text{ erg} \\ m_e &\equiv 9.109 \cdot 10^{-31} \cdot \text{kg} & \frac{h \cdot c}{1 \cdot \text{eV}} &\equiv 1.241 \mu\text{m} \\ \text{Hz} &\equiv \text{sec}^{-1} & \text{Jy} &\equiv 10^{-26} \cdot \frac{\text{W}}{\text{m}^2 \cdot \text{Hz}} & B_v(v, T) &\equiv \frac{2 \cdot h \cdot v^3}{2} \cdot \frac{1}{(h \cdot v)} \end{aligned}$$

Planck