## Handout 15: Virial Theorem E = P.E. + K.E = (1/2)P.E. = -K.E.

- The virial theorem is crucial for an overview of the stellar interior
  - □ See discussion Chap. 2 of C&O
- Holds when P.E. is from a force  $\sim 1/r^2$ 
  - □ i.e. gravity and electrostatic forces
- Consider orbit of satellite around Earth

$$\frac{\mathbf{m} \cdot \mathbf{v}^2}{\mathbf{r}} = \mathbf{F}_{\text{gravity}} = \frac{\mathbf{G} \cdot \mathbf{m} \cdot \mathbf{M}}{\frac{2}{\mathbf{r}^2}}$$
$$\mathbf{KE} = \frac{1}{2} \cdot \mathbf{m} \cdot \mathbf{v}^2 = \frac{1}{2} \cdot \frac{\mathbf{G} \cdot \mathbf{m} \cdot \mathbf{M}}{\mathbf{r}} = \frac{-1}{2} \cdot \mathbf{PE}.$$

## Drag causes increase in speed

- Non-intuitive consequences
  - □ Atmospheric drag causes r to decrease
  - □ P.E. decreases
  - □ |P.E.| increases
  - □K.E. increases
  - Satellite speeds up!
    - Gas drag causes speed to increase
    - Half the decrease in P.E. goes into heating up the atmosphere

## Virial theorem in globular clusters, clusters of galaxies

- Assuming equilibrium has been reached
  - □ High K.E. → high total mass
  - Missing mass discovered in galaxy clusters
    - Zwicky, 1930's
- For the stellar interior, a virial theorem results if
  - $\Box$  Hydrostatic equilibrium dP/dr = -g $\rho$
  - Pressure from ideal gas E.O.S P = nkT
    - And <K.E.> = (3/2)kT
  - □ Massive particles, non-relativistic speeds

#### Derivation of Virial Theorem for stars

$$\frac{d}{dr}P = -\rho \cdot \frac{G \cdot M_r}{\frac{2}{r}}$$

multiply by  $4\pi r^3$  and integrate from r = 0 to R Right Hand Side  $RHS = -\int_{0}^{R} \rho \cdot \frac{G \cdot M_r}{r} \cdot 4 \cdot \pi \cdot r^2 dr$   $4 \cdot \pi \cdot r^2 \cdot \rho \cdot dr = dm$  $-dm \cdot \frac{G \cdot M_r}{r} = dPE$ 

RHS = Gravitational\_PE\_of\_star

#### The virial theorem for stars

LHS = 
$$\int_{0}^{R} \frac{d}{dr} P \cdot 4 \cdot \pi \cdot r^{3} dr$$
  
integrate by parts 
$$\frac{d}{dr} \left( P \cdot 4 \cdot \pi \cdot r^{3} \right) = \frac{d}{dr} P \cdot 4 \cdot \pi \cdot r^{3} + P \cdot 12 \cdot \pi \cdot r^{2}$$
$$P(r = R) = 0 \qquad \text{and} \qquad r(r = 0) = 0 \qquad \text{so}$$

LHS = 
$$-3 \cdot \int_0^K \mathbf{P} \cdot 4 \cdot \pi \cdot \mathbf{r}^2 \, d\mathbf{r} = -3 \cdot \int_0^K \mathbf{n} \cdot \mathbf{k} \cdot \mathbf{T} \cdot 4 \cdot \pi \cdot \mathbf{r}^2 \, d\mathbf{r} = -3 \cdot \frac{2}{3} \cdot \int_0^K \mathbf{n} \cdot \frac{3}{2} \cdot \mathbf{k} \cdot \mathbf{T} \cdot 4 \cdot \pi \cdot \mathbf{r}^2 \, d\mathbf{r} = -2 \cdot KE$$

 $KE = \frac{-1}{2} \cdot PE$  the virial theorem

#### Consequences for stellar evolution

Consider collapse of a star
 As R decreases, PE decreases

$$PE = -\int_{0}^{R} \frac{G \cdot M_{r}}{r} \cdot (4 \cdot \pi \cdot r^{2} \cdot \rho) dr$$

$$PE \text{ will be of order} \qquad \frac{-G \cdot M^{2}}{R}$$
i.e. for constant density
$$PE = \frac{-3}{5} \cdot \frac{G \cdot M^{2}}{R}$$

For a star, gas is compressible, therefore density increases as r decreases Therefore more tightly bound than constant density Therefore PE is more negative than this

## Kelvin-Helmholtz contraction

#### Contract from 100 to 1 R<sub>sun</sub>

- $\square$  PE goes from near zero to  $-GM_{sun}^2/R_{sun}$ 
  - KE increases by  $-(1/2)\Delta PE$ 
    - Star gets hotter inside
  - Where did the other half of the PE go?
    - It was radiated away as photon luminosity
- Kelvin-Helmholtz timescale = -P.E./2\*L

• For the sun at  $R = R_{sun}$ , timescale = 10<sup>7</sup> yr

This is

- Timescale for contraction to the Main Sequence
- Age of the sun if only gravitational energy available

#### Nuclear energy sources

Life on earth is at least 3 to 4 billion years old

- □ Age of sun much longer than 10 million years
- Nuclear energy sources crucial

#### Thermal equilibrium

- Energy flux radiated from the surface exactly equal to the energy generated by nuclear fusion reactions in the interior
- This condition defines the locus of the Main Sequence on the H-R diagram
- $\Box \epsilon(\rho,T)$  is the rate of energy generation, **power/mass**

Thermal equilibrium

$$L_{\text{radiant}} = \int_{0}^{R} 4 \cdot \pi \cdot r^{2} \cdot \rho \cdot \varepsilon \, dr_{\dots}$$

#### Temperature in the sun

The virial theorem gives the average temperature of the sun

$$\frac{3}{2} \cdot k \cdot T \cdot N = \frac{-1}{2} \cdot PE = G \cdot \frac{M_{sun}^2}{R_{sun}}$$

$$N = \frac{M_{sun}}{\mu \cdot m_H} \qquad \mu = 0.6$$

$$T_{ave} = \frac{2}{3} \cdot \frac{\mu \cdot m_H \cdot G}{k} \cdot \left(\frac{M_{sun}}{R_{sun}}\right) = 10^7 \cdot K \qquad \text{amazing}$$

Just by looking at the sun for  $\sim$  1 hour  $\rightarrow$  hydrostatic equilibrium  $\rightarrow$  Virial theorem  $\rightarrow$  interior is 10 million K!

# Sun temperatures of 10 million K → nuclear fusion reactions

- High velocities necessary to overcome coulomb barrier
- 4 H  $\rightarrow$  eventually, He + energy
  - □ And energy ~ 1% mc<sup>2</sup>
    - Very significant, c.f. (1/2)PE
      - $\Box$  1% Mc<sup>2</sup> = 10<sup>52</sup> ergs
      - $\Box GM^{2}/2R = 10^{48} ergs$
    - Rather than the K-H time of 10<sup>7</sup> yr, we have 10<sup>11</sup> years for nuclear, Main Sequence lifetime

□ Stellar evolution  $\rightarrow$  only 10% of H burnt on M.S.